Are the laws of physics actually fine-tuned?

Jim Clarage Department of Chemistry and Physics University of St. Thomas

G

Gravitational Force

Gravity mysteries: Why is gravity fine-tuned?

"The feebleness of gravity is something we should be grateful for. If it were a tiny bit stronger, none of us would be here to scoff at its puny nature."

1 part in 10¹⁵

0.00000000000001



proton and neutron

 $m_p = 1.6726 \times 10^{-27} \text{ kg}$ $m_n = 1.6749 \times 10^{-27} \text{ kg}$

If $m_p > m_n$ proton decays to neutron (no atoms)If $m_p << m_n$ deuteron unstable (no pp reaction)If $m_p \sim = m_n$ no beta decay possible

proton-proton chain reaction



 $4^{1}H \rightarrow {}^{4}He + 2e^{+} + 2\gamma + 2v_{e} (26.7 \text{ MeV})$

Implications...

- Anthropic Principle
- Ø Designer (aka Tuner)
- Multiverse

Anthropic Principle (Brandon Carter, 1974)

Weak:

What we can expect to observe must be restricted by the condition necessary for our presence as observers.

Strong:

The Universe (and hence the fundamental parameters on which it depends) must be such as to admit the creation of observers within it at some stage.

Hoyle's Carbon resonance

triple-alpha process

 $3 {}^{4}\text{He} \rightarrow {}^{12}\text{C}$

prediction: 7.65 MeV (1952) measured: 7.656 MeV

Critique of fine-tuning

Literature



Le baron chassa Camilide du châncan. (Page 3.)

CANDIDE

L'OPTIMISME.

CHAPITRE PREMIER.

Comment Candide fut élevé dans un beau château, et comment il fut chase d'icelui-

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Voltaire's Candide, 1759

1

"There is a concatenation of all events in the best of possible worlds..."

-Master Pangloss, teacher of the metaphysico-theologo-cosmolonigology

History...





History...



William Paley (1743–1805) used the watchmaker analogy in his book Natural Theology, or Evidences of the Existence and Attributes of the Deity collected from the Appearances of Nature, published in 1802.

April, 1931

THE NATURE OF THE CHEMICAL BOND

[CONTRIBUTION FROM GATES CHEMICAL LABORATORY, CALIFORNIA INSTITUTE OF TECHNOLOGY, NO. 280]

THE NATURE OF THE CHEMICAL BOND. APPLICATION OF RESULTS OBTAINED FROM THE QUANTUM MECHANICS AND FROM A THEORY OF PARAMAGNETIC SUSCEPTIBILITY TO THE STRUCTURE OF MOLECULES

BY LINUS PAULING

RECEIVED FEBRUARY 17, 1931 PUBLISHED APRIL 6, 1931

During the last four years the problem of the nature of the chemical bond has been attacked by theoretical physicists, especially Heitler and London, by the application of the quantum mechanics. This work has led to an approximate theoretical calculation of the energy of formation and of other properties of very simple molecules, such as H2, and has also provided a formal justification of the rules set up in 1916 by G. N. Lewis for his electron-pair bond. In the following paper it will be shown that many more results of chemical significance can be obtained from the quantum mechanical equations, permitting the formulation of an extensive and powerful set of rules for the electron-pair bond supplementing those of Lewis. These rules provide information regarding the relative strengths of bonds formed by different atoms, the angles between bonds, free rotation or lack of free rotation about bond axes, the relation between the quantum numbers of bonding electrons and the number and spatial arrangement of the bonds, etc. A complete theory of the magnetic moments of molecules and complex ions is also developed, and it is shown that for many compounds involving elements of the transition groups this theory together with the rules for electron-pair bonds leads to a unique assignment of

Goldilocks planet

not too hot (nor cold)

oxygen

sunlight

"Has the Earth a Corner on Life?" Kenneth Crist, Los Angeles Times, 1935

GUILLERMO GONZALEZ AND JAY W. RICHARDS

AUTHORS REVIEWS Documentary Q&A

LINKS

The PRIVILEGED PLANET

Read a synopsis of the book

order now

HOW OUR PLACE IN THE COSMOS IS DESIGNED FOR DISCOVERY

QuickTime[™] and a Sorenson Video 3 decompressor are needed to see this picture.

Goldilocks planet

not too hot (nor cold) thermophilic (118°C !)

- oxygen anaerobic (no oxygen)
- sunlight heat, sulfur



Aquifex pyrophilus

How to un-fry and egg





Cystine

Apollo 12 (1969)



Surveyor 3's camera (1967)



Streptococcus mitis

Language

"delicately dependent..."
"slightly different..." (14)
"very slightly stronger..."
"wildly improbable numerical accidents..."
"somewhat weaker..."
"tiny bit stronger..."

"Slightly" stats: 14 times in Davies, "The Accidental Universe" 20 times in Barrow, "The Constants of Nature: The Numbers That Encode the Deepest Secrets of the Universe"



$F_1 = 979.677941576350$ $F_2 = 979.677634377790$

 $\Delta F \sim 1/10^{7}$

m = 100 kg, dr = 1.0m

Rare isotopes in the cosmos

Hendrik Schatz

Such stellar processes as heavy-element formation and x-ray bursts are governed by unstable nuclear isotopes that challenge theorists and experimentalists alike.

Hendrik Schatz is a professor of physics and astronomy at Michigan State University's department of physics and astronomy and National Superconducting Cyclotron Laboratory in East Lansing. He is also cofounder and associate director of the Joint Institute for Nuclear Astrophysics, an NSF Physics Frontier Center.

Radioactive nuclei with extreme neutron-to-proton ratios—rare isotopes—often decay within fractions of a second. Typically, they are not found on Earth unless produced at an accelerator. Yet nature produces copious amounts of them in supernovae and other stellar explosions in which the rare isotopes, despite their fleeting existence and small scale, imprint their properties. Large amounts of them also exist as stable layers in the crusts of neutron stars.

Rare isotopes are therefore intimately linked to fundamental questions in astrophysics. An example is the origin of the 50-odd naturally occurring elements between the iron region and uranium in the periodic table. As figure 1 shows, with recent progress in stellar spectroscopy and the continuing discovery of very old and chemically primitive stars, a "fossil record" of chemical evolution is now emerging. Nuclear science needs to make its own progress to match specific events to the observed elemental abundance patterns produced in stellar explosions. That has turned out to be a tremendous challenge. Nuclear theory is still far from being Nuclear Astrophysics in the US exemplifies such a center. Across the Atlantic, Europe is witnessing such initiatives as the Extreme Matter Institute and the Munich Cluster of Excellence on the Origin and Structure of the Universe, both in Germany, and the international Challenges and Advanced Research in Nuclear Astrophysics network.

Origin of the elements

Nuclear processes in stars and stellar explosions have forged nature's chemical elements out of the hydrogen and helium left over from the Big Bang. Fusion reactions in stars drive the formation of elements with atomic numbers up to about that of iron. But fusion does not produce heavier elements; the nuclear binding energy per nucleon is maximal for nuclei around iron and nickel, so continued fusion would be endothermic.

Heavier elements are thought to be built up by a neutron-capture process that iterates a two-step sequence. First, neutrons bombard a seed nucleus until a number have been

feature article

Hoyle's Carbon resonance

triple-alpha process

 $3 {}^{4}\text{He} \rightarrow {}^{12}\text{C}$

Steven Weinberg: fine = 20%

Chemistry

Carbon fixation





Mathematics

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4	19 20 1 K Ca 39.098 40.08 4	21 Sc 44.956	22 Ti 47.88	23 V 50.942	24 Cr 51.996	25 Mn 54.938	26 Fe 55.847	27 Co 58 933	28 Ni 58.69	29 Cu 63.546	30 Zn 65.39	31 Ga 69.72	32 Ge 72.59	33 As 74.922	34 Se 78.96	35 Br 79.904	36 Kr 83.80
5	37 38 3 Rb Sr 85.468 87.62 8	39 Y 88.906	40 Zr 91.224	41 Nb 92.906	42 Mo 95.94	43 Tc (98)	44 Ru 101.07	45 Rh 102.91	46 Pd 106.42	47 Ag 107.87	48 Cd 112.41	49 In 114.82	50 Sn 118.71	51 Sb 121.75	52 Te 127.60	53 126.91	54 Xe 131.29
6	55 56 Cs Ba 132.91 137.33	57 to 71	72 Hf 178.49	73 Ta 180.95	74 W 183.85	75 Re 186.21	76 Os 190.2	77 Ir 192.22	78 Pt 195.08	79 Au 196.97	80 Hg 200.59	81 TI 204.38	82 Pb 207.2	83 Bi 208.98	84 Po (209)	85 At (210)	86 Rn (222)
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Biology

Stating the obvious

 r_a = radius axle holes r_b = radius lug bolt

If $r_a > r_b$ unstable, decays 1-2mins.

If $r_a < r_b$ bound state cannot form!

Conclusion: travel as we know it impossible w/out tuning.

Why don't biologists have an Anthropic Principle?

Biology





Cytochrome *c*,

Computer Science

Life qua Algorithm



Life qua Algorithm



If C(t+1) >= C(t) {
TumbleTime = long;
}
Else {
TumbleTime = short;

Architecture Independent



Systems Biology of Human Aging - Network Model 2009

Rev. 9 June 2009 © 2000 - 2009 John D. Furber All rights reserved.

Arrangement, text, & art by John D. Furber

www.LegendaryPharma.com/chartbg.html



Complex Systems

Pre-stating

"Never calculate anything until you already know the answer"

-Landau



TATA-box binding protein (TBP)

(c) ETH Zurich & MedILS

MOLECULAR MACHINERY: A Tour of the **Protein Data** Bank

会にな 3cyt Cytochrome c 1gen Glucagon 2hlu Insulin **3hhr Human Growth** Hormone 1rfb Interferen 2daj Degryribonuclease

1121 Lysoryme **2ptc** Trypsin **5rsa** Ribonucleose Spep Papsia 1smd Amylose

6.07 **1poe** Phospholipose **ligt** Antibody

4etry Rhinovirus

1c17+1e79 ATP Synthese 1bgy Cytechrome bel Complex 1000 Cytochrome c Oxidate 1/88 Rhadepsin 1pth Cyclosxygenese **1prc** Photosynthetic **Reaction Center** 1kzu Light-Harvesting Complex

Cholesterol

Phospholipid

85

1|b0 Photosystem |

Synthetese



1fqy Aqueporin



4tna Tronsfer

DMA

Pre-stating, non-ergodicity

Time required to create all possible 200 residue proteins at least once:

10⁶⁷ times the lifetime universe!

Stuart Kauffman, Investigations



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