#### Gullies on Mars -- Water or Not?







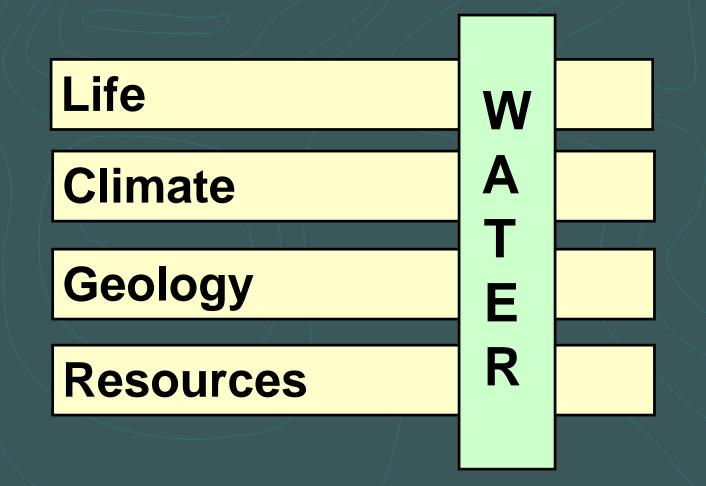
#### Diam. 12,742 km g=9.80 m/s<sup>2</sup>

Diam. 6,780 km Diam. 3,474 km g=3.71 m/s<sup>2</sup> g=1.62 m/s<sup>2</sup>

Allan H. Treiman



## NASA Mars Exploration Strategy: "Follow the Water"

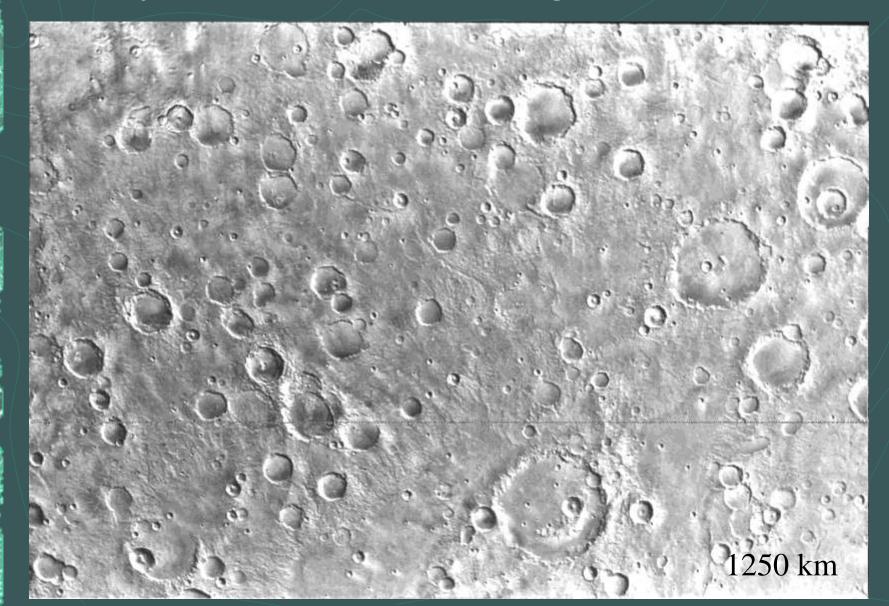


## Evidence of Water on Mars

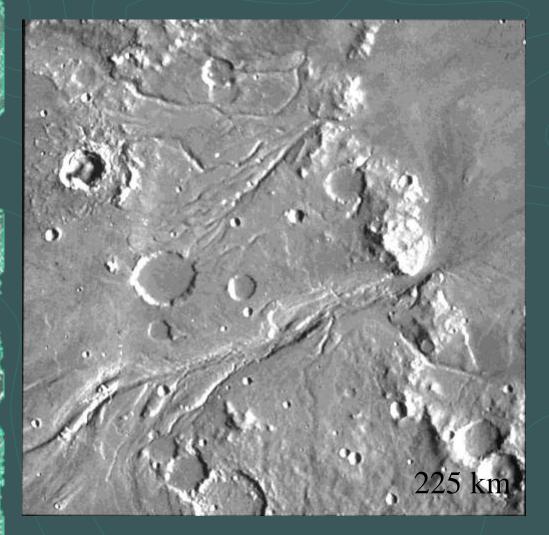
#### Distant Past

- Crater Degradation and Valley Networks
- 'River' Channels
- Flat Northern Lowlands
- Meteorites
  - Carbonate in ALH84001
  - Clay in nakhlites
- MER Rover Sites Discoveries
  - Hydrous minerals: jarosite!
  - Fe<sub>2</sub>O<sub>3</sub> from water (blueberries etc.)
  - Silica & sulfate & phosphate deposits
- Recent Past (Any liquid?)
  - Clouds & Polar Ice
  - Ground Ice

## Valley Networks and Degraded Craters



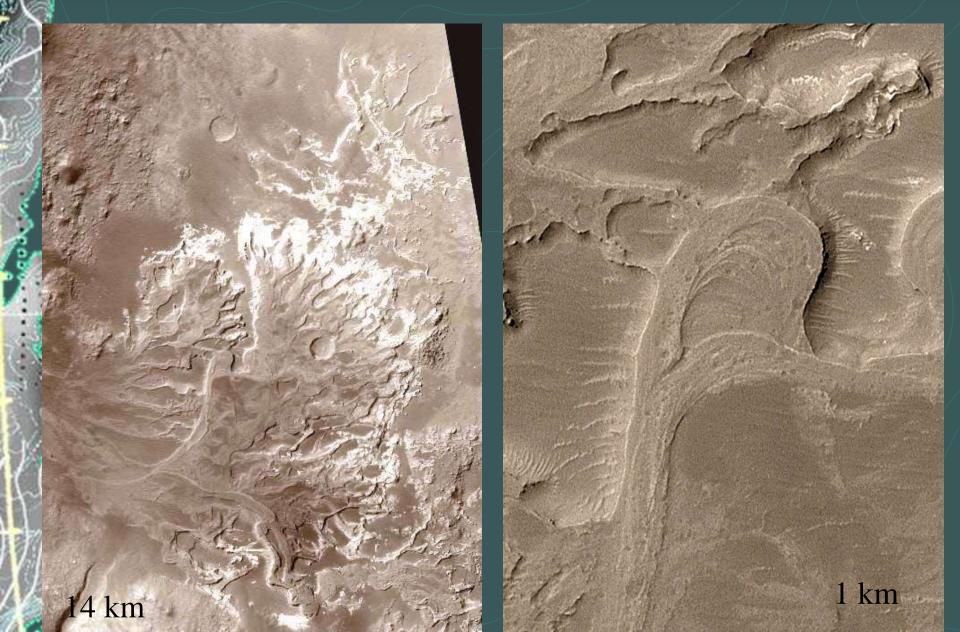
## **River Channels - Giant Floods!**





10 km craters

## River Channels - 'Normal' Flows



# River Channels from Rain?

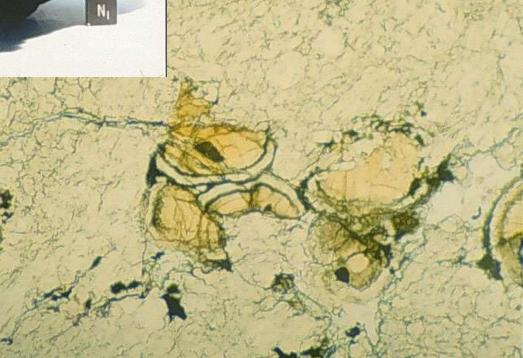


Science, July 2, 2004



## Ancient Martian Meteorite ALH84001

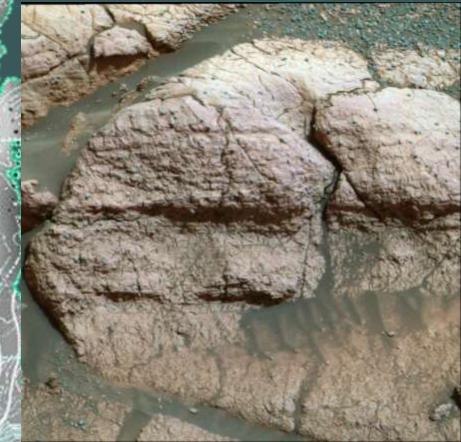




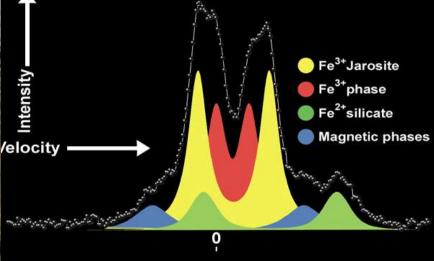
# MER Opportunity - Heatshield and parachute.



Jarosite - A Water-bearing Mineral Formed in Groundwater KFe<sub>3</sub><sup>3+</sup>(SO<sub>4</sub>)<sub>2</sub>(OH)<sub>6</sub> 2 Jarosite = K<sub>2</sub>SO<sub>4</sub> + 3 Fe<sub>2</sub>O<sub>3</sub> + 3 H<sub>2</sub>SO<sub>4</sub>





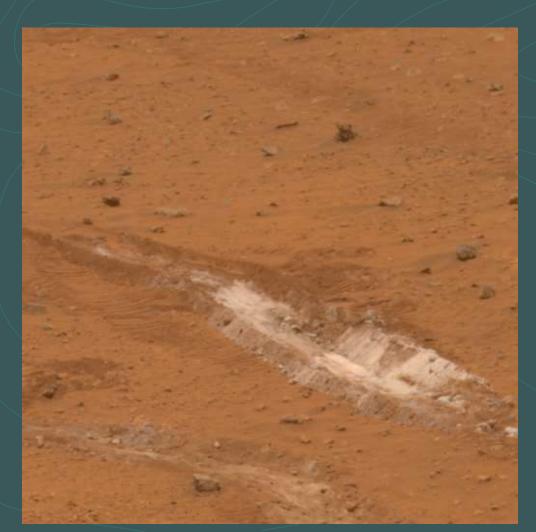


## Hematite is in "Blueberries," which still suggest water.

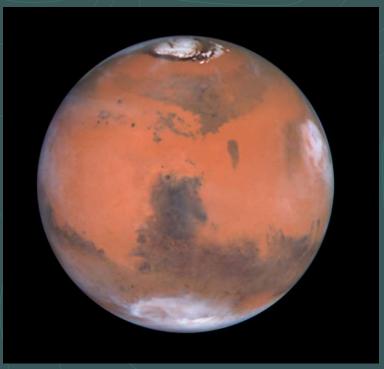
#### **Stone Mountain**



## MER Spirit: Columbia Hills



## H<sub>2</sub>O Now: Clouds & Polar Caps

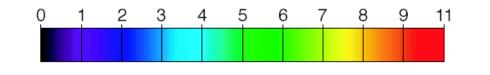


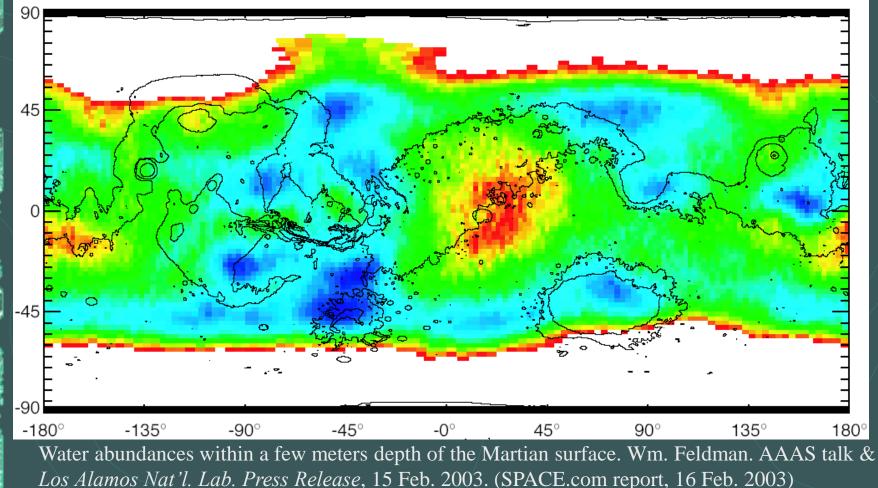




## Ground Ice – Mars Orbiter GRS

#### Water Abundance (weight percent)





### So, Water on Mars !!

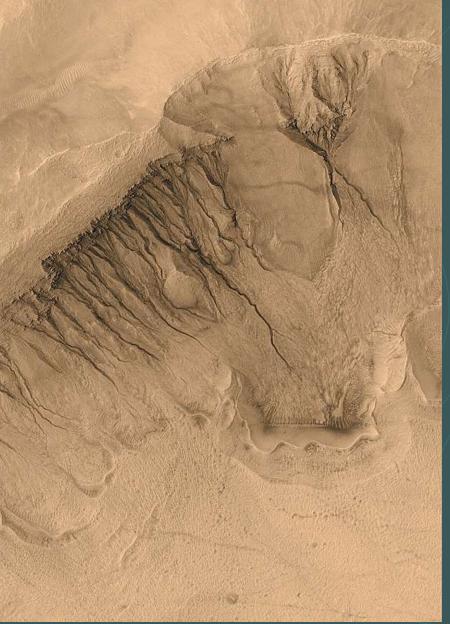
So? Apparently, Mars has/had lots of water.
Lots of evidence for ancient liquid water (> ~2 billion years ago), both surface and underground.

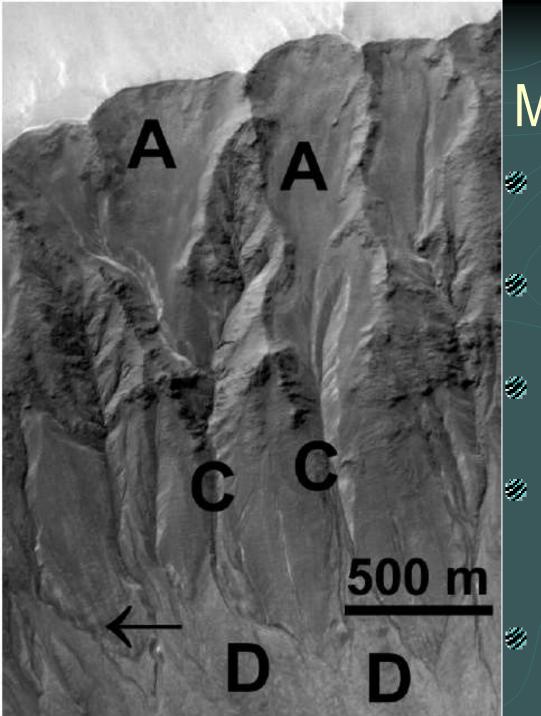


## Martian Gullies - Liquid Water or Not?

 Material flows down steep slopes, most commonly interpreted as water-bearing debris flows [Malin and Edgett (2000) Science 288, 2330].

 Liquid water is difficult to produce and maintain near Mars' surface, now.
Could water be stable (enough) recently at Mars' surface?





#### Mars Gully Mass flow on steep slopes Alcoves - sources of material Channels - many sinuous, with levees Deposits - cones of debris, common with distinct lobate toes. Best analogs - water-rich debris flows

## Most Earth analogs involve liquid water



Problems making liquid water Too cold! Except  $10^{3}$ occasionally 10<sup>4</sup> (bars) Heat source Critical Liquid Solid Gas 10 point Antifreeze Pressure Pressure too low? Triple 10 point Atmospheric Antifreeze pressure /apor 10 Form liquid at higher 10 pressure, eject to surface 0 Solid Gas 10 Form liquid rapidly at surface, and make flow 100 200 300 400 500 600 700 before it freezes/boils Temperature (K) away.

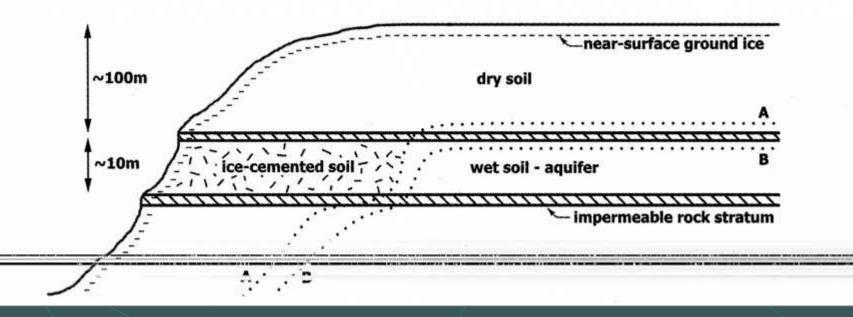
800

## Internal Liquid Water - Groundwater.

- Sunlight heats proper slopes (mid-latitude polefacing) just right (Malin & Edgett)
- Only enough solar heating when Mars' spin axis tilted more than today (Costard et al.)
- Internal heat is enough insulated by deep fluffy dust (Mellon) [implies regional aquifers]
- Heat from shallow magma (Hartmann)
- Water erupts from depth (Gaidos)
- No heat needed, water is briny (Burt et al.)

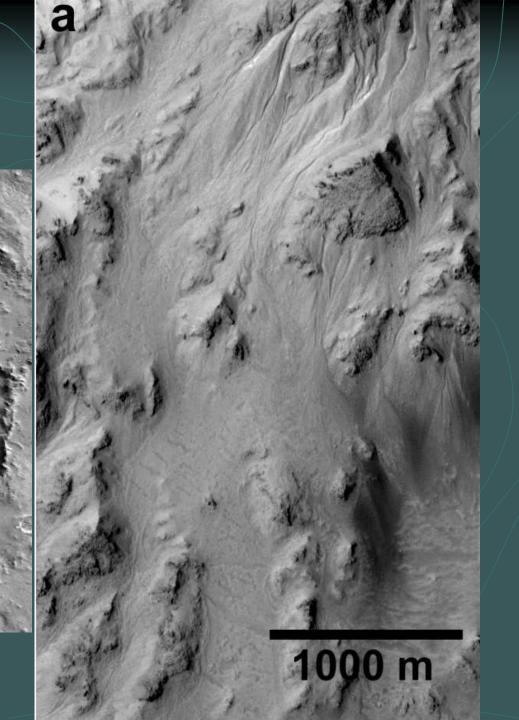
## Pressurized (Burst Pipe)

- Liquid water trapped, bounded by impermeable layers of rock, ice, frozen soil.
- In cooling climate or annual cycle, water begins to freeze.
- Volume increase on freezing pressurizes liquid.
- Excess pressure breaks rock. Whoosh!

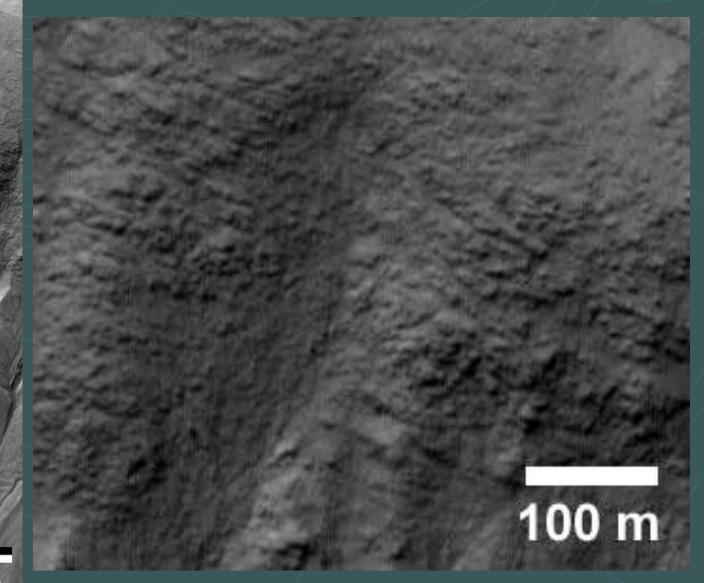


## Impermeable Layers ?

Central Peak, Hale impact basin



## Crater wall: Newton area

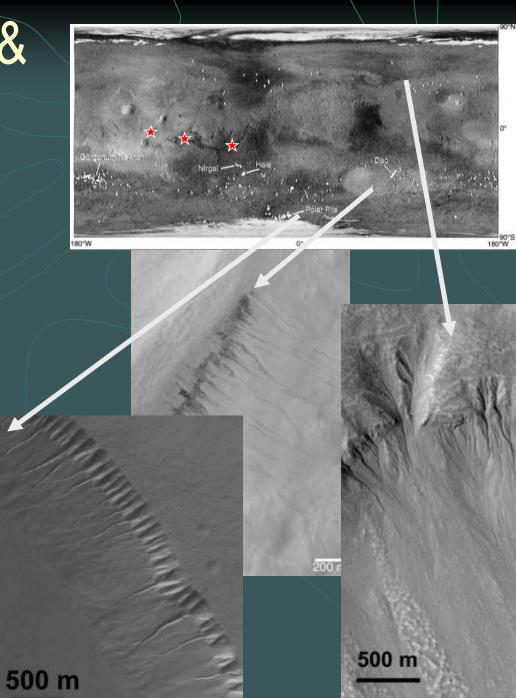




### Liquid Water from Snowmelt. Christensen (2003) Wind-blown snow accumulates on pole-facing slopes (i.e., mantling deposits). BUT, snow accumulates in lee of obstacles, I.e. on equator-facing slopes. Top anneals to 'impermeable' layer. Ø Own weight provides a few extra millibars pressure. Sunlight into snow heats it to melting. Gullies form beneath snowcover!

## Most Altitudes & Latitudes

Altitude Pavonis summit Hellas floor Latitude S. peri-polar (75S) Equatorial N. peri-polar (69N) Altitude/Latitude e.g., Pavonis / Capri Hale / Hellas



## Ground Ice Melting?

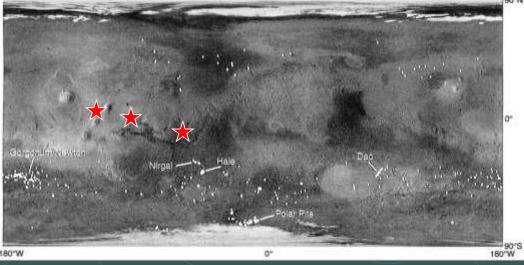
- Ground ice preserved from past climate epoch (higher obliquity)
- Sunlight warms slopes, causes melting & fluidization of debris.
  - Quicker than the melt water can evaporate/boil away.

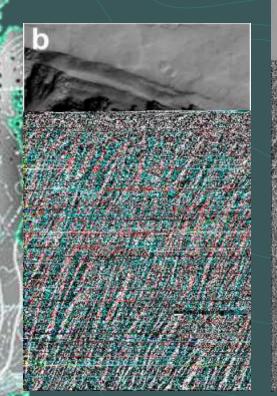
Wind can remove dust and expose fresh ice to produce additional flows.

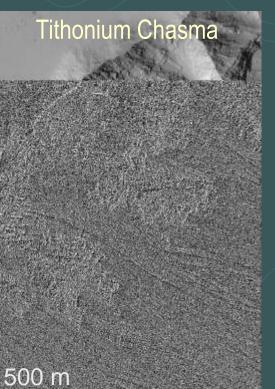
## Near-Equatorial

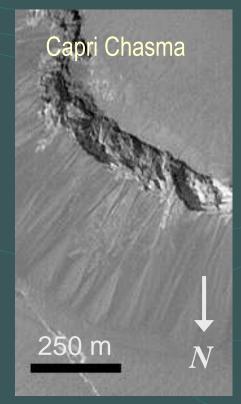
Summit Caldera, Pavonis Mons

## Gullies









#### Trouble all around. Solar heating on slopes? (Obliquity, snowmelt) 4 Gullies on slopes of all orientations. Gullies at all latitudes and elevations. Pressurized aquifers? (Water, brine) Most gullies on broken and/or unlayered rock. How could groundwater get to tops of isolated peaks?



## Why consider 'dry' mechanisms?

No liquid today at Mars' surface
Liquid water barely stable near Mars' surface
Skepticism about generation of liquid water
Pressurizing aquifers in broken rock of crater walls?

- Bringing water to mountain-tops?
- Just right' melting of ground ice / snow?
- Good Earth Analogs
  - Sand' etc.
  - Snow
  - Pyroclastic

## Analogs: Sand

Click here to view the video

## Analog: Snow

Advantages Ø Dry snow is abundant Forms very large flows Abundant in steep slopes of large vertical relief Extensive study - avalanche safety Disadvantages Rheology complicated in detail Can partly melt in large flows

## Snow as a Fluid



## Sinuous Channels: Sionne & Gotthard

## Analog: Pyroclastics Mt. St. Helens 1980





Figure 300. — Perspective view of a single lobe of deposits from July 22 eruption showing slightly raised levees and central channel. Persons are standing at edge of channel just inside levees. Dashed line marks approximate crest of levees. Lobe turns to right just behind persons. Photograph taken at station B (fig. 298).

## Wilson & Head (1981) USGS. PP 1250

The Argument: Morphology  $\Rightarrow$  Rheology  $\Rightarrow$  Composition "Morphology  $\Rightarrow$  Rheology" is dicey Must understand <u>all</u> physics completely! E.g., lava flow inflation not recognized early "Rheology Composition" is not unique Gully levees, sharp-front flows suggest Bingham flow law Consistent with water-'granule' slurries, gas-supported

or dry granules, some lavas, some emulsions, etc.

Dry / Gas-supported Flows are Legitimate Alternatives. Advantages & Disadvantages

- Require nothing unobserved (only slopes, grains, gas)
  Extensive theoretical, experimental bases
- No exact analogs on Earth, not near so much dry dust / sand.
- Theory, experiments only at much smaller length scales

## More to come!

MRO/HiRISE

Imagery **@ THEMIS** MiRISE Topo. #IRISE Spectra CRISM

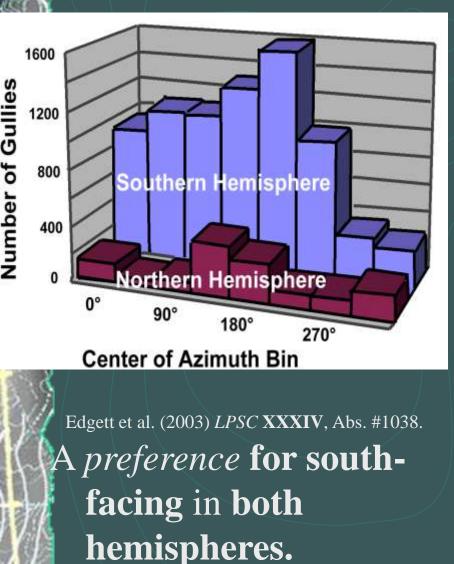
2.5 km

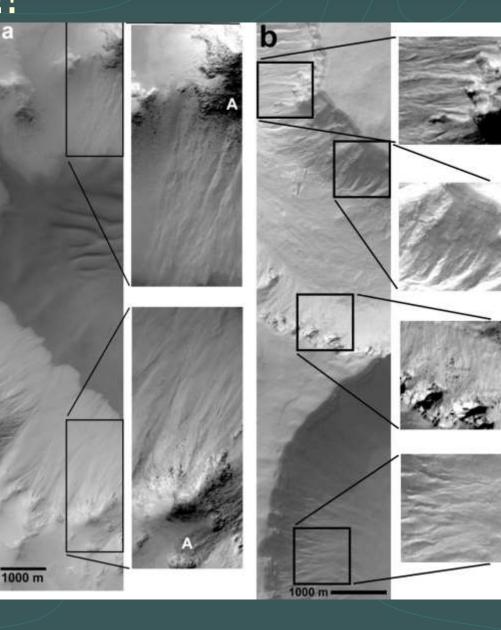
THEMIS

500 meters

PSP\_006648\_1300

## Not Pole-facing !!





**Problems and Unknowns** Regional / global distribution of gullies Consistent with dry(±gas) flows? Ø Different in the past? Local distributions of gullies Consistent with dry(±gas) flows? Theory / Experiments Are results at meter scale relevant? Can dry(±gas) flows really meander?

### Argument by Analogy

"… this word analogy is urged, as giving great force to the reasoning. But it must be recollected, that precisely the point in question is whether there is an analogy." Whewell W. (1853) Of a Plurality of Worlds: An Essay. London.

… is risky at best.